Question		ion	Answer	Marks	Guidance
1	(a	(i)	$H_{0} = 1/\text{age}$ $H_{0} = 1/(13.7 \times 10^{9} \times 3.16 \times 10^{7})$ $(H_{0} =) 2.31 \times 10^{-18} \text{ (s}^{-1})$ $(H_{0} =) \frac{2.31 \times 10^{-18} \times 3.09 \times 10^{16} \times 10^{6}}{10^{3}}$ Hubble constant = 71.4 (km s ⁻¹ Mpc ⁻¹)	C1 C1 A1	Allow: 2 sf answer Special case : Using $H_0 = 1/13.7 \times 10^9 = 7.30 \times 10^{-11} (y^{-1})$
		(ii)	$v = H_0 d$ (v =) 71.4 × 50 or 3.57 × 10 ³ (km s ⁻¹) or 3.57 × 10 ⁶ (m s ⁻¹)	C1	gives an answer of 2.26×10^9 (km s ⁻¹ Mpc ⁻¹) – allow 1 mark Possible ecf from (a)
			$\frac{\Delta\lambda}{\lambda} = \frac{3.57 \times 10^6}{3.0 \times 10^8} (= 1.19 \times 10^{-2})$	C1	
			$\Delta \lambda = 656 \times 1.19 \times 10^{-2}$ or $\Delta \lambda = 7.80$ (nm) wavelength = 656 + 7.80	C1	
			wavelength = 664 (nm)	A1	Allow: 2sf answer
	(b)		Big bang: Creation of the universe (from which space/time evolved) (AW) Any <u>three</u> from:	B1	
			 (At the start) the universe was hot / infinitely dense Expansion of the universe led to cooling The (current) temperature of universe is 2.7 K / 3 K (The universe as a black body) is associated with microwaves at this temperature (AW) or The (wavelength of the) gamma radiation stretched to microwaves (by the expansion). 	B1 × 3	Not : The universe now has microwaves. (The microwaves must be linked with current temperature)
			QWC: (Cosmological principle is supported because) MBR is isotropic	B1	Allow: Microwaves have the same intensity in all directions

Question	Answer	Marks	Guidance
(c)	(For an open / flat universe)		
	Further expansion will lead to cooling / temperature lower than 3K / temperature tend to absolute zero (AW)	B1	Alternative:Temperature (will eventually) increases if closed universe B1The wavelength (of EM radiation) get smallerB1
	The wavelength (of the EM radiation) gets longer / frequency (of the EM radiation) gets smaller / energy of photons decreases / microwaves become radio waves	B1	
(d)	Graph starting from origin and having a shape consistent with either open or accelerated universe	B1	Not a straight line
	Total	15	

Q	uestio	Answer	Marks	Guidance
2	(a)	 Any <u>four</u> from: 1. (Fusion is the) joining / fusing together of ('lighter') <u>nuclei</u> / <u>protons</u> (to make 'heavier' nuclei) 2. Mass decreases in the reaction and this is transformed into energy OR the products have greater binding energy 3. High temperatures / ~10⁷ K needed for fusion 4. High pressure / density (required in the core) 5. The protons / nuclei repel (each other because of their positive charge) 6. The strong (nuclear) force comes into play when the protons / nuclei are close to each other 	B1×4	Not: Atoms / particles for nuclei /protons.
	(b)	 (When hydrogen / helium runs out) the outer layers of the star expands / a (super) red giant is formed The core (of the star) collapses (rapidly) / a <u>supernova</u> is formed (Depending on the initial mass of the star the remnant is either a) <u>neutron star</u> or a <u>black hole</u> 	B1 B1 B1	
		Total	7	

Question		on	Answer	Marks	Guidance
3	(a)		$F = \frac{GMm}{r^2}$ force = $\frac{6.67 \times 10^{-11} \times (10^{41})^2}{(4 \times 10^{22})^2}$	C1 C1	
			$(4 \times 10^{22})^2$ force = 4.2 × 10 ²⁶ (N)	A1	Allow: 4×10^{26} (N) or 10^{26} since this is an estimation Allow: 2 marks for 4.2×10^{n} ; n $\neq 26$ (POT error)
	(b)		 Allow any <u>one</u> from: The galaxies are receding / moving away from each other (because of the big bang) Other galaxies may be pulling them in opposite direction The acceleration is too small to collapse (other than over a very long period of time) 	B1	
	(c)		 Any <u>six</u> from: 1. (At the start it was) very hot / extremely dense / singularity 2. All forces were unified 3. Expansion led to cooling 4. Quarks / leptons (soup) 5. More matter than antimatter 6. Quarks combine to form hadrons / protons / neutrons 7. Imbalance of neutrons and protons / (primordial) helium produced 8. Atoms formed 9. Idea of gravitational force responsible for formation of stars / galaxies 10. Temperature becomes 2.7 K / 3 K or (the universe is saturated with cosmic) microwave background radiation 	B1×6	Show annotation on Scoris
	(d)	(i)	Dark lines / bands against a background of <u>continuous</u> <u>spectrum</u>	M1 A1	

Question	Answer	Marks	Guidance
(ii)	$\frac{v}{c} = \frac{\Delta\lambda}{\lambda}$ speed = $\frac{86.6}{393.4} \times 3.0 \times 10^8$ (Any subject) speed = 6.6×10^7 (m s ⁻¹) or 66000 (km s ⁻¹) $v = H_0 d$ $66000 = 50 \times d$ distance = 1300 (Mpc)	C1 C1 A1	Allow: 1 mark for $\frac{86.6}{480.0} \times 3.0 \times 10^8 = 5.41 \times 10^7$ (m s ⁻¹) Allow: 2 marks for 1.3×10^n ; n \neq 3 (POT error) Note: Answer is 1080 (Mpc) if 5.4×10^7 (m s ⁻¹) is used; this value will score 2 marks
	Total	15	

Question	Expected Answers	Marks	Additional guidance	
4 (a)	The critical density is the density for which the universe will expand towards a (finite) limit or rate of expansion tends to zero / which will result in a <u>flat</u> universe	B1	Not: critical density is given by $\frac{3{H_0}^2}{8\pi G}$	
(b)	Hubble constant = $\frac{65 \times 10^{3}}{10^{6} \times 3.1 \times 10^{16}}$ Hubble constant = 2.1 × 10 ⁻¹⁸ s ⁻¹ critical density = $\frac{3H_{0}^{2}}{8\pi G}$ critical density = $\frac{3 \times (2.1 \times 10^{-18})^{2}}{8\pi \times 6.67 \times 10^{-11}}$ critical density = 7.9 × 10 ⁻²⁷ (kg m ⁻³)	B1 C1 A1	Possible e.c.f. from value of Hubble constant within this calculation	
(c) (i)	open: (density of universe < critical density hence) the universe will expand forever closed: (density of universe > critical density hence) the universe will (eventually stop expanding and then) contract / big crunch flat: (density of universe = critical density hence) the universe will expand towards a (finite) limit / rate of expansion tends to zero	B1 B1 B1	 Allow: 'universe continues to expand' Not: 'The universe stops expanding' Special case: Award 1 mark for correct sketches if no explanation is given for open, closed and flat 	
(ii	Any <u>one</u> from: Existence of dark matter / black holes / neutrinos / dark energy / H ₀ is not known accurately	B1		
	Total	8		